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Complex Behaviour in Classical and Quantum Chaos

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DYNAMICAL SYSTEMS AND COMPLEXITY

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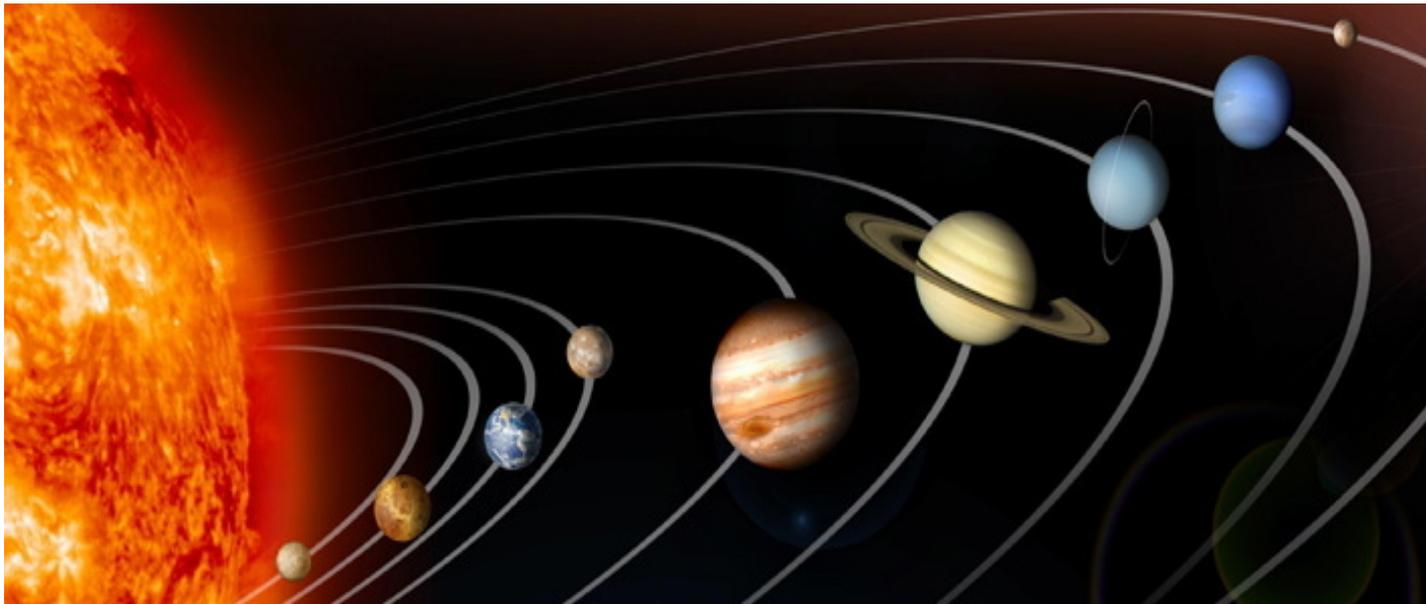
Calandra, Halkidiki, 28 August - 6 September 2024

ABSTRACT

I shall explain how chaos (chaotic behaviour) can emerge in deterministic systems of classical dynamics. It is due to the sensitive dependence on initial conditions, meaning that two nearby initial states of a system develop in time such that their positions (states) separate very fast (exponentially) in time. After a finite time (Lyapunov time) the accuracy of orbit characterizing the state of the system is entirely lost, the system could be in any allowed state. The system can be also ergodic, meaning that one single chaotic orbit describing the evolution of the system visits any other neighbourhood of all other states of the system.

In this sense, chaotic behaviour in time evolution does not exist in quantum mechanics. However, if we look at the structural and statistical properties of the quantum system, we do find clear analogies and relationships with the structures of the corresponding classical systems. This is manifested in the eigenstates and energy spectra of various quantum systems (mesoscopic solid state systems, molecules, atoms, nuclei, elementary particles) and other wave systems (electromagnetic, acoustic, elastic, seismic, water surface waves etc), which are observed in nature and in the experiments.

The Solar System of 8 (or 9) planets (out of scale)



Mercury, Venus, Earth, Mars, Jupiter, Saturn, Uranus, Neptune, (Pluto)

On the short time scale: the trajectories are ellipses (Kepler)

On the long run, the ellipses can stretch or shrink, rotate and tilt, even in a chaotic way.

The divergence/separation of nearby orbits in regular and chaotic systems are drastically different:

Slow: Linear in regular systems: Separation $\propto t$

Fast: Exponential in chaotic systems: Separation $\propto \exp(\gamma t)$

Lyapunov exponent = γ , and Lyapunov time = $\frac{1}{\gamma}$

For times much larger than Lyapunov time the motion is unpredictable

Example: In the case of Pluto: Lyapunov time \approx 20 million years

Wisdom and Susskind (1988) and Jacques Laskar (since 1990)

On the long run for certain initial conditions the planets and other bodies like asteroids and comets might move chaotically and collide with each other, or escape from the Solar System.

Nevertheless, in the case of Earth, we know that the motion is stable and regular, on an almost circular orbit, over about 5 billion years. This is crucial for the stability of climate and existence of life.

Motivation by the example of a simple dynamical system

Two-dimensional classical billiards:

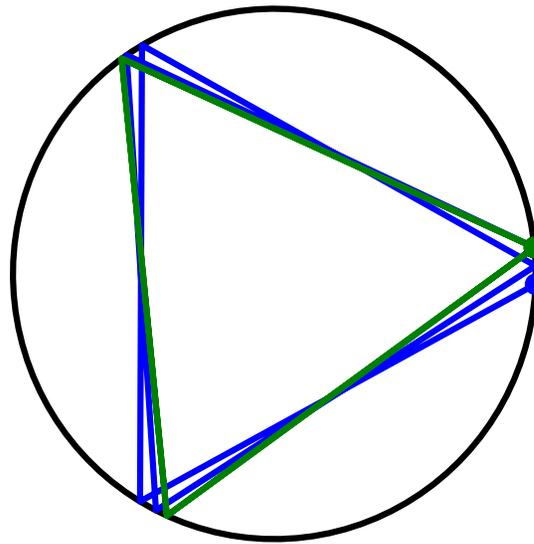
A point particle is moving freely inside a two-dimensional domain with specular reflection on the boundary upon the collision:

Energy (and the speed) of the particle is conserved.

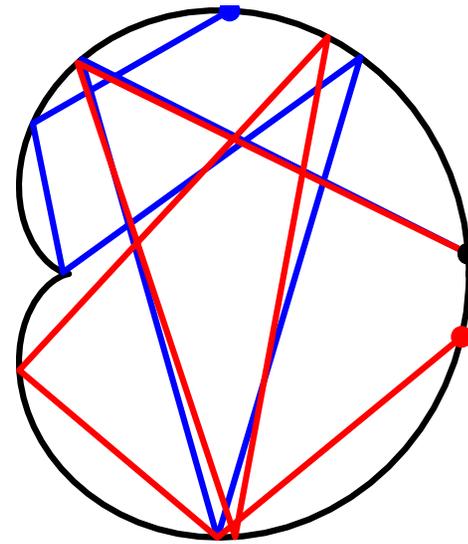
A particular example of the billiard boundary shape as a model system:

Complex map: $z \rightarrow w, |z| = 1$

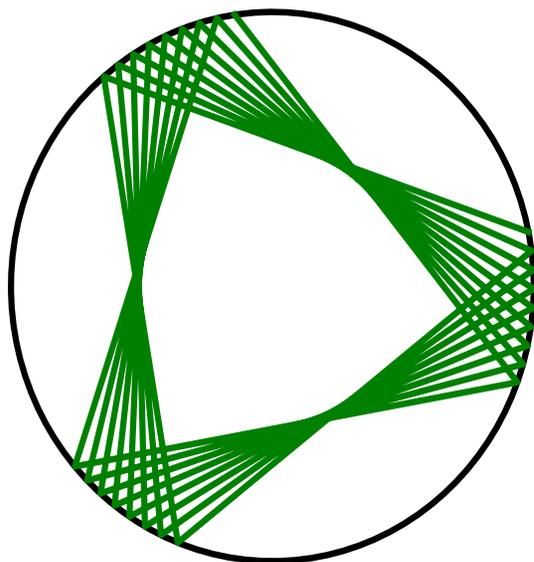
$$w = z + \lambda z^2,$$



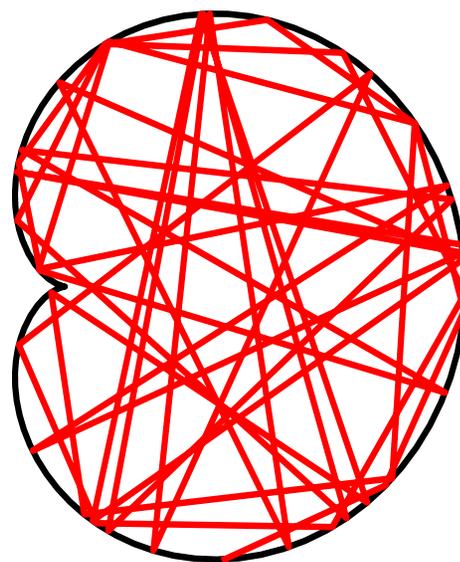
$\lambda = 0$



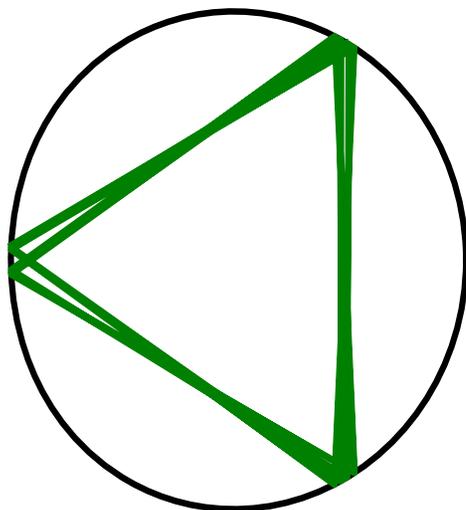
$\lambda = 0.5$



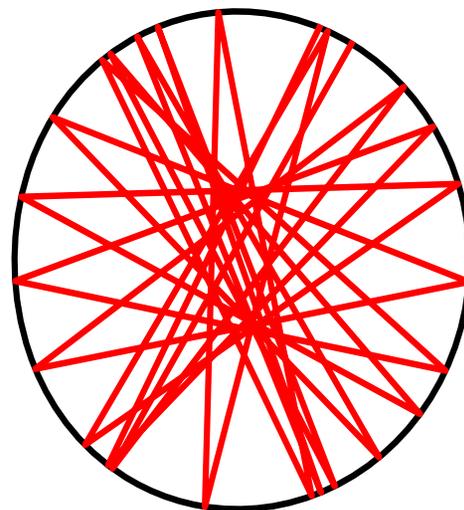
$\lambda = 0$



$\lambda = 0.5$



$\lambda = 0.15$



QUANTUM CHAOS: Motivation by the simple example

Two-dimensional quantum billiards: a point particle trapped in a two-dimensional box

Helmholtz (Schrödinger) equation with Dirichlet boundary conditions

$$\frac{\partial^2 \psi}{\partial x^2} + \frac{\partial^2 \psi}{\partial y^2} + E\psi = 0$$

with $\psi = 0$ on the boundary

Solutions exist only at the discrete values of E : the eigenenergies.

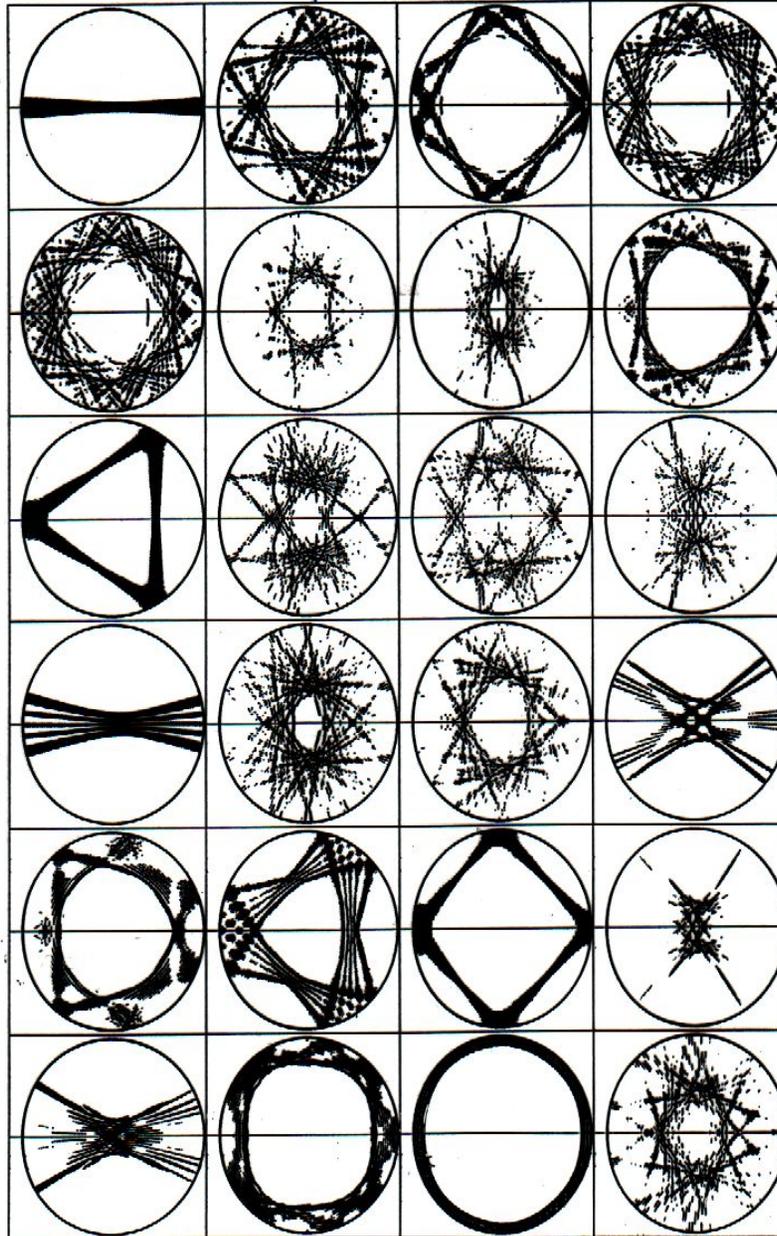
They are infinitely many, but countable.

*This equation also describes the oscillations of an elastic membrane (a drum):
 E is the square of the eigenfrequency of the drum. ψ is the amplitude.*

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$W = z + \lambda z^2$
Figure 1 (Robnik 1983)

$\lambda = 0.15$



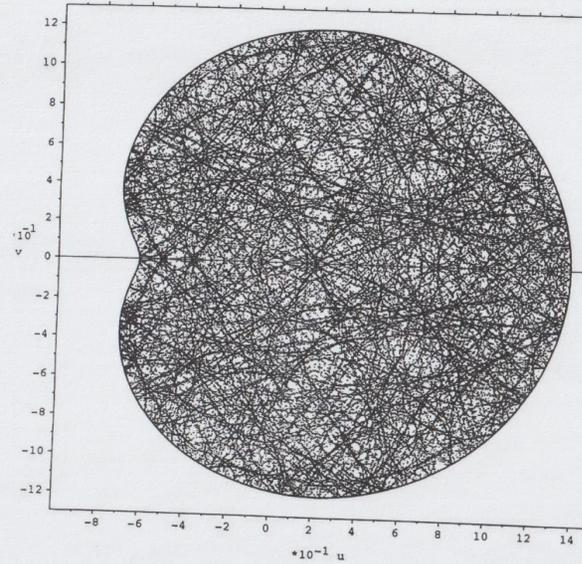
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Figure 3

$$\lambda = 0.375$$

Figure 4

$$W = \vartheta + \lambda \vartheta^2$$



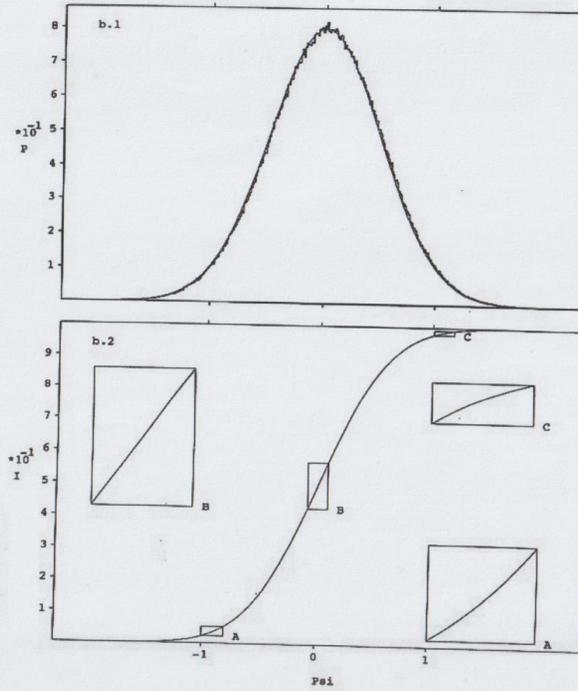
126 de Broglie wavelengths

even parity $n \approx 100\,015$ $E = 625\,118.4$
even

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Figure 2b

Figur 10



Statistical properties of discrete energy spectra with the same density

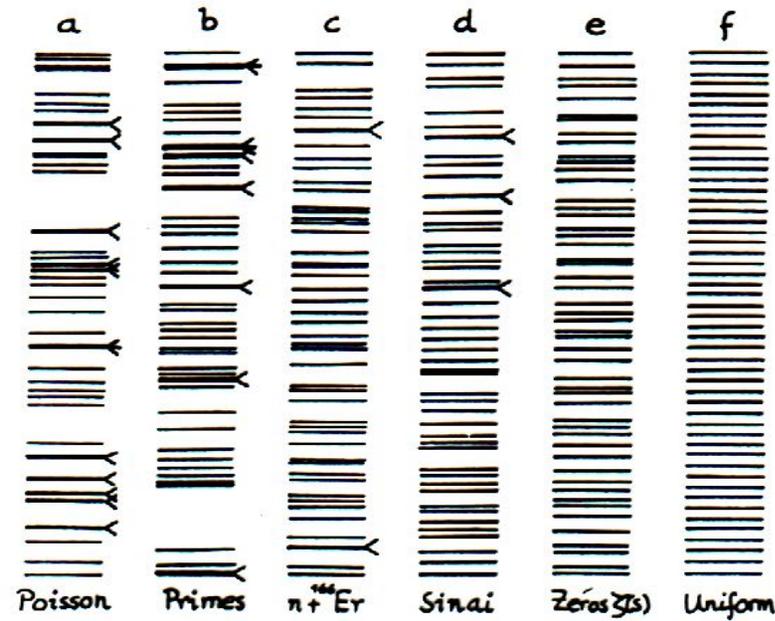
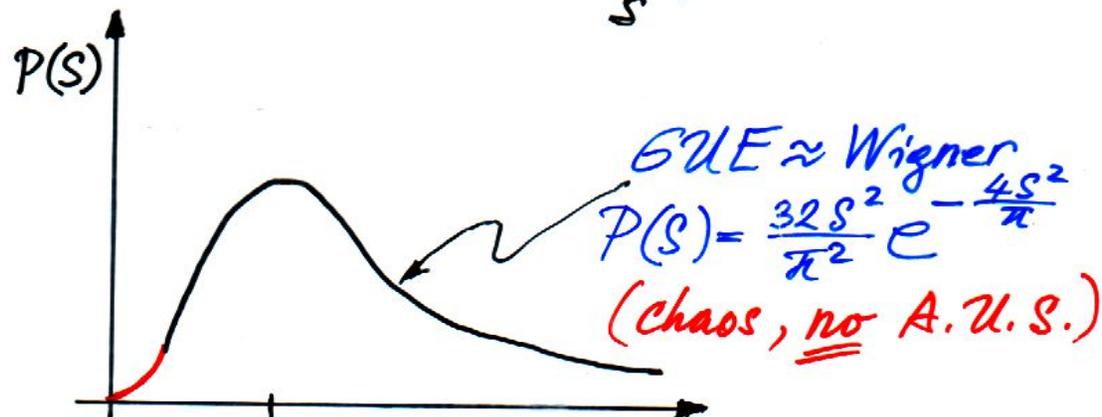
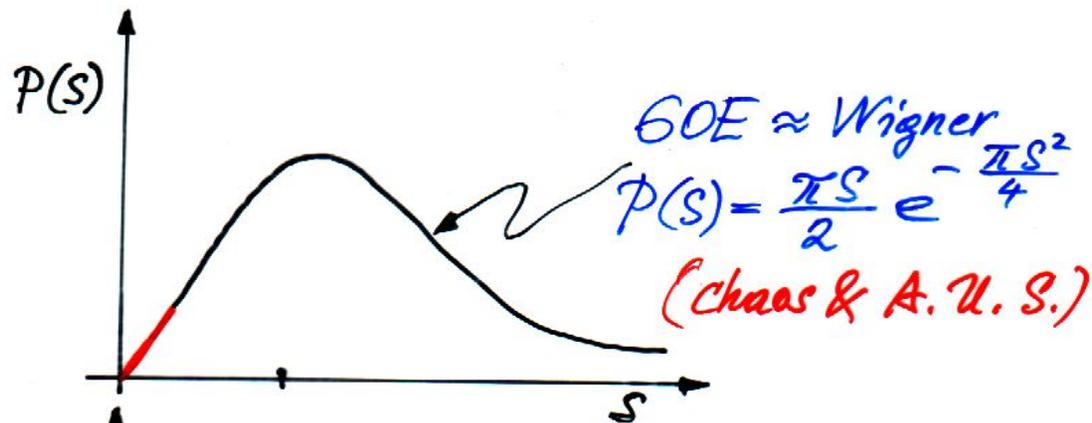
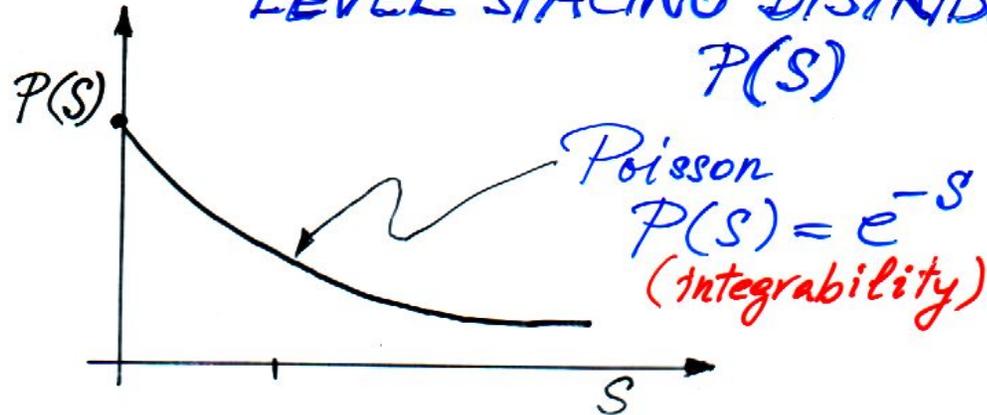


Fig.1.8 - Segments of "spectra", each containing 50 levels. The "arrowheads" mark the occurrence of pairs of levels with spacings smaller than $1/4$. See text for further explanation.

Bohigas and Giannoni 1984

ENERGY (NEAREST NEIGHBOUR)
LEVEL SPACING DISTRIBUTION



If the system is of the mixed type, partially regular and partially chaotic, thus when regular and chaotic regions coexist in the classical phase space, we have semiclassical theories, which describe the intermediate region between the Poissonian and GOE/GUE statistics. They are well established and follow the so-called Principle of Uniform Semiclassical Condensation of Wigner functions in the quantum phase space (PUSC).

But let us look again at the classical mixed-type systems

(2)

integrable Hamiltonian systems:

N integrals (constants) of motion exist
 N = number of degrees of freedom

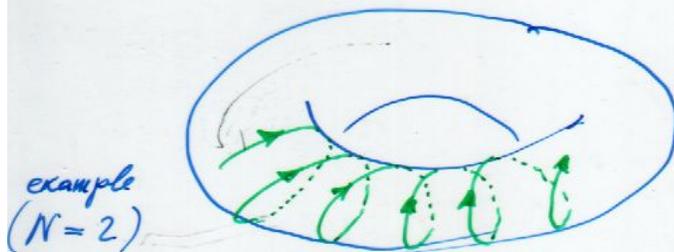
$$A_i = A_i(\vec{q}, \vec{p}) = A_i(\vec{q}(t), \vec{p}(t)) = \text{const.}$$

$$i = 1, 2, \dots, N \quad (A_1 = E = H(\vec{q}, \vec{p}))$$

$$\{A_i, A_j\} = \text{Poisson bracket} = 0, \forall i, j$$

$$= \frac{\partial A_i}{\partial \vec{q}} \cdot \frac{\partial A_j}{\partial \vec{p}} - \frac{\partial A_i}{\partial \vec{p}} \cdot \frac{\partial A_j}{\partial \vec{q}} = 0$$

Liouville-Arnold theorem:



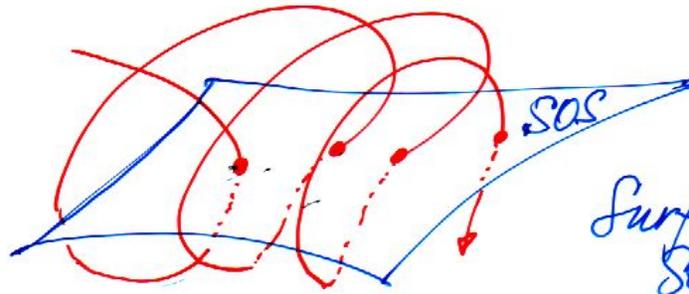
N -dim
 invariant
 tori
 (for all initial
 conditions)

The ergodic systems (fully chaotic):

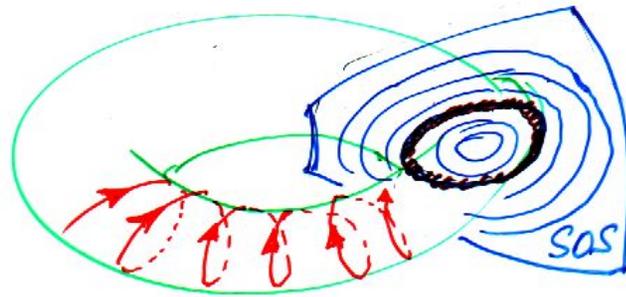
no integrals of motion except
 the total energy $E = H(\vec{q}, \vec{p}) = \text{const.}$

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(3)



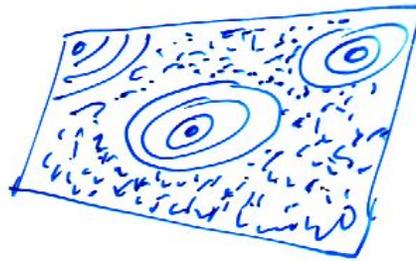
Surface of Section



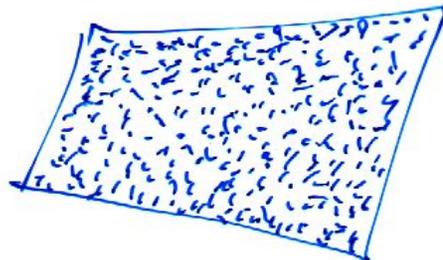
integrable:
N-tori
everywhere

perturb ↓

Kolmogorov
Arnold
Moser



KAM
picture



ergodic
and
chaotic

Example of mixed type system: Hydrogen atom in strong magnetic field

$$H = \frac{\mathbf{p}^2}{2m_e} - \frac{e^2}{r} + \frac{eL_z}{2m_e c} |\mathbf{B}| + \frac{e^2 \mathbf{B}^2}{8m_e c^2} \rho^2$$

\mathbf{B} = magnetic field strength vector pointing in z -direction

$r = \sqrt{x^2 + y^2 + z^2}$ = spherical radius, $\rho = \sqrt{x^2 + y^2}$ = axial radius

L_z = z -component of angular momentum = conserved quantity

Characteristic field strength: $B_0 = \frac{m_e^2 e^3 c}{\hbar^2} = 2.35 \times 10^9$ Gauss = 2.35×10^5 Tesla

Rough qualitative criterion for global chaos: magnetic force \approx Coulomb force

(Wunner et al 1978+; Wintgen et al 1987+; Hasegawa, R. and Wunner 1989, Friedrich and Wintgen 1989; classical and quantum chaos: R. 1980+)

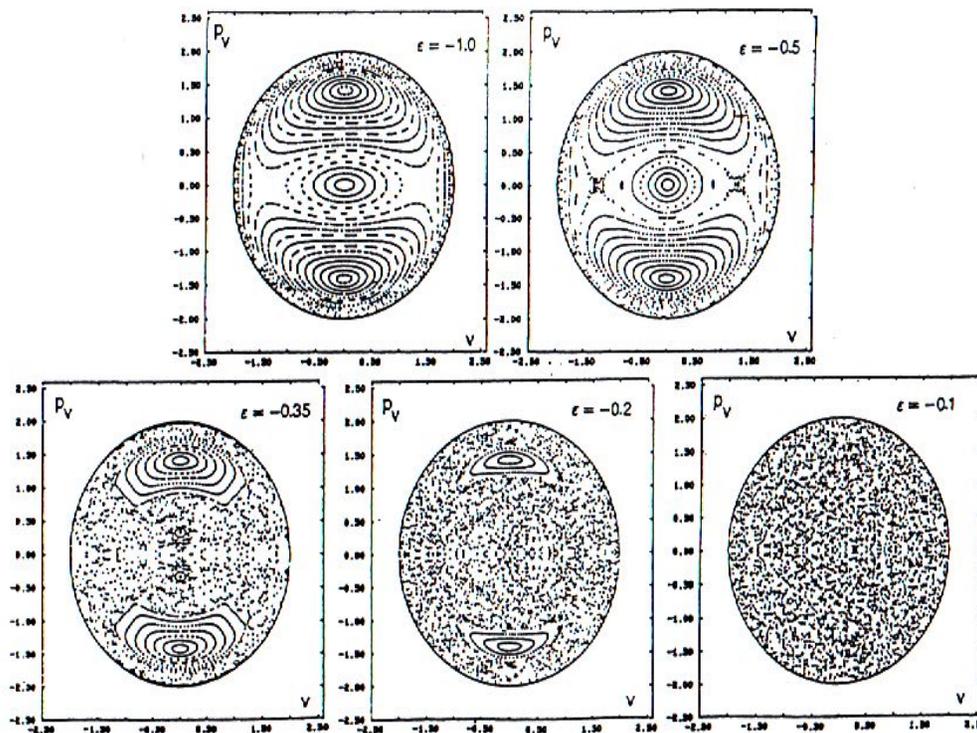


Fig. III-9. Poincaré surfaces of section $\Sigma(v, p_v; u=0)$ at different scaled energies (corresponding to increasing diamagnetic strength). The elliptic fixed point at the origin corresponds to the straight-line orbit I_0 , the other two fixed points to the straight-line orbit I_1 .

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Thank you very much!